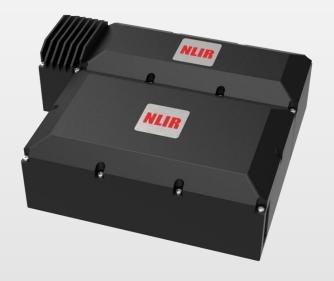


## **Mid-Infrared Light Detector**

Fast - Sensitive - Rugged



- Single-wavelength detection from 2.2 – 5.0 μm
- DC 10 GHz bandwidth
- NEP down to 2 fW/ $\sqrt{\text{Hz}}$
- Based on novel upconversion technology

## **NLIR** Mid-Infrared Single-Wavelength Detector

- a new paradigm in mid-infrared light detection

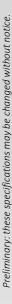
Measuring mid-infrared (MIR) light is inherently difficult due to the small photon energies and finite temperatures of detectors, which results in low detection efficiencies and high levels of noise. The NLIR MIR light detector is based on a novel measurement scheme that upconverts the MIR light to the near-visible regime. Near-visible light detectors (based on for example Si) are far superior to MIR light detectors in terms of efficiency, speed and noise. The NLIR upconversion technology therefore brings these attractive features and the advantages that follow to the MIR regime.

New regimes of the MIR region can be explored using the NLIR technology: with up to 10 GHz sampling rate, nanosecond pulses are possible to characterise directly in the time domain or, alternatively, light at specific wavelengths from chemical reactions can be measured at unprecedented much faster than possible otherwise. New regimes of sensitivity also come within reach: the noise-equivalent power of the NLIR single-wavelength detector is as low as 2 fW/ $\sqrt{\rm Hz}$ , which more than an order of magnitude lower than state-of-the-art cooled HgCdTe (MCT) and InSb MIR detectors.

## Key applications:

- Nanosecond pulse characterisation
- Low-power MIR light detection
- · Detection of gas lines

- Free-space communication
- DIAL (differential absorption LIDAR)
- Chemical kinetics





## **Mid-Infrared Light Detector**

Fast - Sensitive - Rugged

The NLIR MIR light detector is composed of an upconversion unit that changes the wavelength of the MIR light to the near-visible regime and a near-visible light detector. Near-visible light detectors exist with many different combinations of specifications and what detector to choose depends on the desired application. Below, two examples give an overview of possible operation specifications:

Example 1: Fast	
Optical bandwidth	< 10 nm
Centre wavelength	3.3 µm
Noise-equivalent power (NEP)	$0.1\mathrm{nw}/\sqrt{\mathrm{Hz}}$
Saturation limit	300 μW
Electrical bandwidth	20 kHz – 10 GHz
Responsivity	130 V/W
Field-of-view	<2°
Detector size	200 μm
Optical input	Free space, polarisation sensitive (incl. guide beam)
Connector type	Male SMA
Physical dimensions ( $h \times l \times w$ )	70 mm × 230 mm × 260 mm

Centre wavelength available from 2200 – 5000 nm. Optical bandwidth and FOV change slightly with the choice of centre wavelength.

Example 2: Sensitive	
Optical bandwidth	< 10 nm
Centre wavelength	3.3 µm
Noise-equivalent power (NEP)	$3 \text{ fw}/\sqrt{\text{Hz}}$
Saturation limit	50 pW
Electrical bandwidth	DC – 10 MHz
Responsivity	30 000 kV/W
Field-of-view	<2°
Detector size	200 μm
Optical input	Free space, polarisation sensitive (incl. guide beam)
Connector type	Male SMC
Physical dimensions ( $\mathbf{h} \times \mathbf{l} \times \mathbf{w}$ )	70 mm × 230 mm × 260 mm

Centre wavelengths are currently available from  $2.5-4.5~\mu m$ . Optical bandwidth and FOV change slightly with the choice of centre wavelength.